Remote Sensing

Remote Sensing Literature:

- Lillesand T.M. and Kiefer R.W., 2000. Remote sensing and image interpretation. John Wiley & Sons.
- Aronoff S., 2005. Remote sensing for GIS managers; ESRI Press, Redlands, California.
- Barrett E.C. and Curtis L.F., 1992. Introduction to environmental remote sensing. Chapman & Hall.
- Jensen J.R., 2000. Remote sensing of the environment. An earth resource perspective. Prentice Hall, Upper Saddle River, New Jersey.

Remote Sensing Table of contents

- Introduction to remote sensing
- History of Remote Sensing
- Multispectral Remote Sensing Systems
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REMOTE SENSING DEFINITIONS:

<u>The more general</u>: Remote Sensing involves gathering data and information about the physical "world" by detecting and measuring signals composed of radiation, particles, and fields emanating from objects located beyond the immediate vicinity of the sensor device(s).

<u>The more restricted</u>: Remote Sensing is a technology for sampling electromagnetic radiation to acquire and interpret non-contiguous *geospatial data* from which to extract information about features, objects, and classes on the Earth's land surface, oceans, and atmosphere



"Remote" and "Proximal" Sensing

Nicholas Short Tutorial of Remote Sensing (http://rst.gsfc.nasa.gov/Intro)

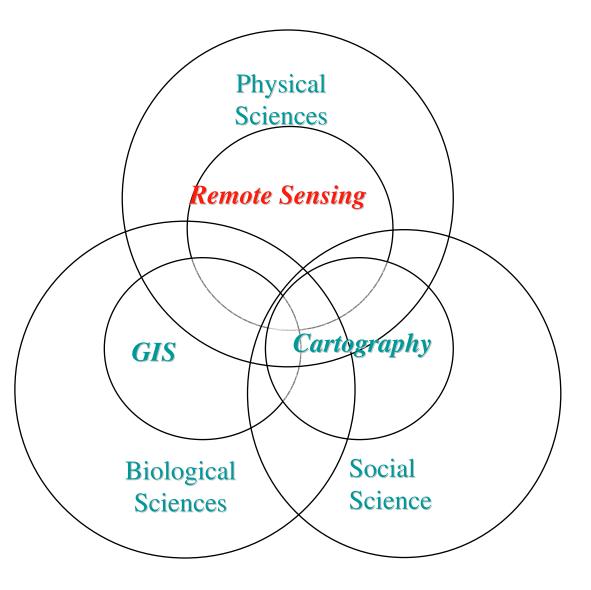
Remote Sensing methods:

- passive analog (camera, videography)
- passive digital (camera, multispectral scanners, linear and area arrays, spectroradiometers)
- active (microwave radar, laser lidar, sonar)

Remote Sensing observations can be carried out from a level:

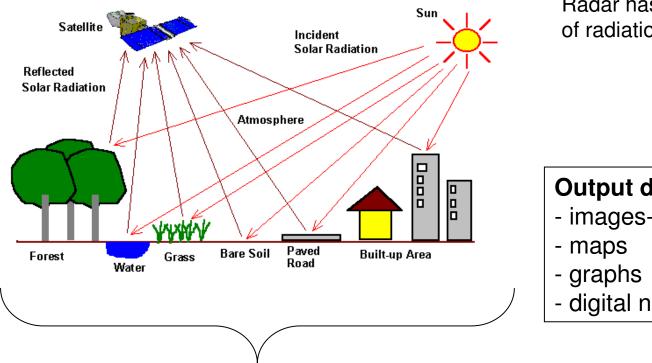
- Satellite,
- Airborne,
- Low Altitude Remote Sensing (LARS)
- Ground.

Relationships between Remote Sensing, Cartography and Geographical Information System on a background of a physical, biological and social sciences



Why remote sensing is also an art?

- The ability of visual photo or digital image interpretation derives not only from scientific knowledge, but also from the experience acquired during a past life.
- Some image analyst are much superior to other image analysts because they:
 - Understand the scientific principles better,
 - are more widely travelled and have seen many landscape objects and geographic areas first-hand
 - can synthetize scientific principles and real-world knoowledge to reach logical and correct conclusions.



Radar has its own source of radiation.

Output data:

- images-photography
- digital numbers.

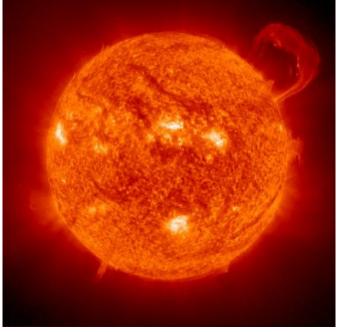
Incident solar radiation can be:

reflected, transsmitted and/or absorbed in various proportions depending on material characteristics

Solar energy – is created in the process of thermosynthesis - in which, by a combination of four protons in the nucleus of helium, the mass conversion into energy occurs.

Energy is the ability to do work Can be transferred in: conduction, convection and radiation

Radiation - transfer of energy that do not require physical contact between the body and absorbing radiant energy. Energetic particles or energetic waves travel through a medium or space. They **radiate** (i.e., travel outward in all directions) from a source.

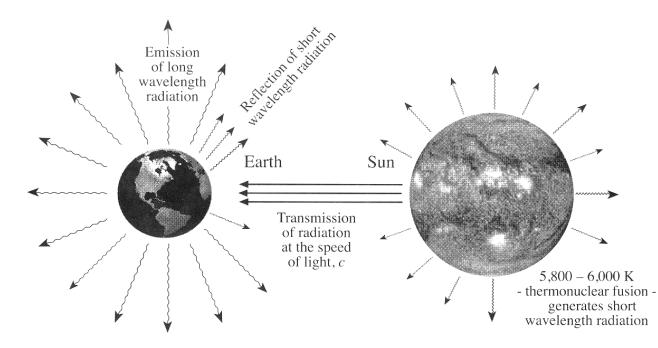


Radiation:

- ionizing (or ionising) radiation is radiation composed of particles that individually carry enough kinetic energy to liberate an electron from an atom or molecule, ionizing it. Ionizing radiation is generated through nuclear reactions, either artificial or natural, by very high temperature

- **non-ionizing**: the light from the Sun that reaches the earth is largely composed of non-ionizing radiation. Far-ultraviolet rays have been filtered out by the gases in the atmosphere, particularly oxygen.

Wave model – James C. Maxwell



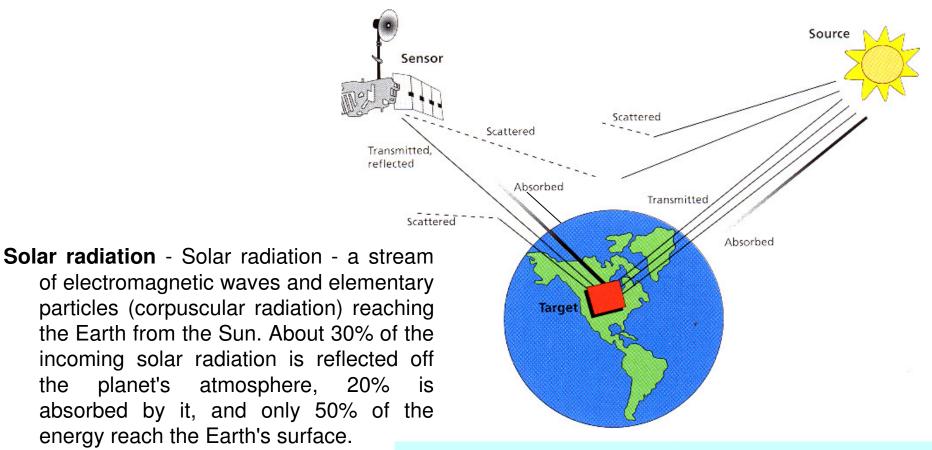
Blackbody – a

theoretical construct that absorbs and radiates energy at the maximum possible rate per unit area at each wavelength for a given temperature

Stefan-Boltzmann law – the total emitted radiation from a blackbody [W m⁻²] is proportional to the fourth power of its absolute temperature.

$$M_{\lambda} = \sigma T^4$$

TELEDETEKCJA podstawy fizyczne



Solar radiation:

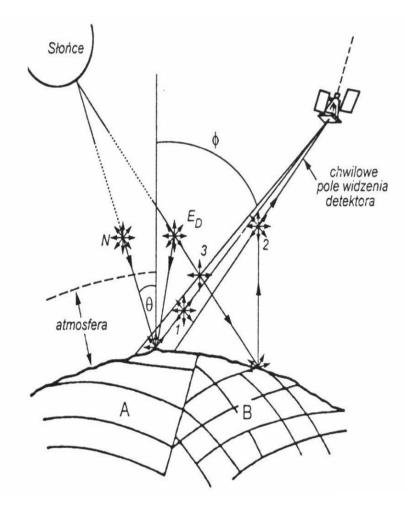
- direct,

- diffuse (sky radiation) solar radiation from the hemisphere of the sky reaching the Earth's surface, with the exception of direct radiation,

- total.

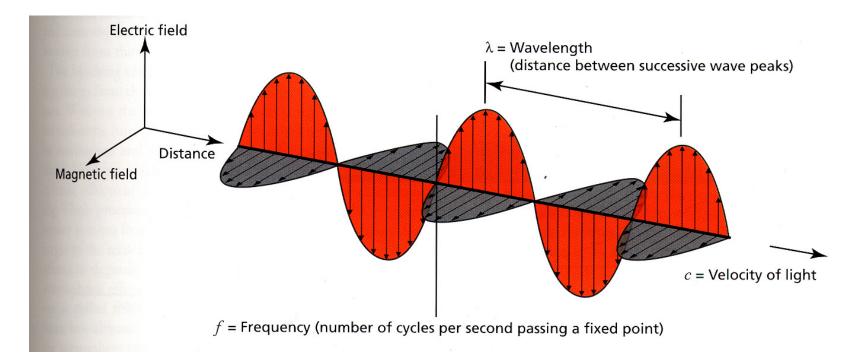
The energy from the Sun being remotely sensed:

- is radiated by atomic particles at the source (the Sun),
- propagates through the vacuum of space at the speed of light,
- interacts with the Earth's atmosphere,
- interacts with the Earth's surface,
- interacts with the Earth's atmosphere once again, and
- finally reaches the remote sensing, where it interacts with various optical systems, filters, film emulsions, or detectors.



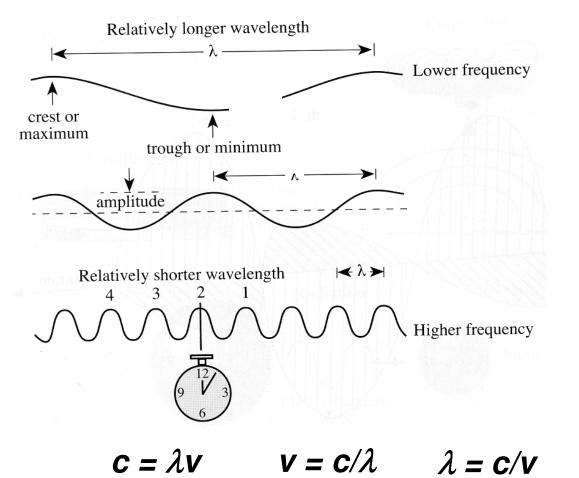
Wave model – James C. Maxwell

An electromagnetic wave is composed of both electric and magnetic vectors that are orthonogal (at 90° angles) to one another



The waves travel from the source at the speed of light (3 x 10⁸ m s⁻¹)

Wave model – James C. Maxwell



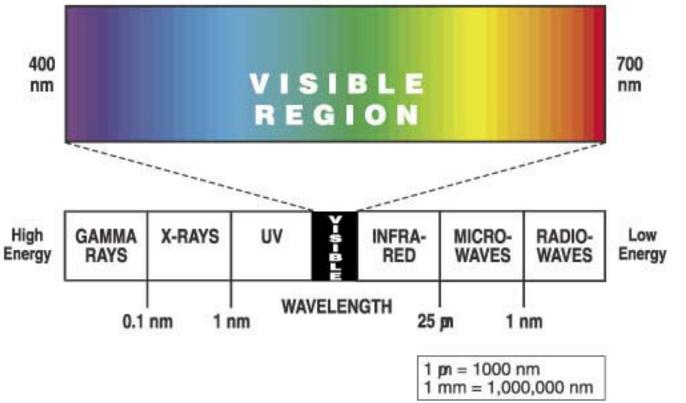
Wavelength - the mean distance between maximums (or minimums) of a roughly periodic pattern [µm or nm].

Frequency - the number of wavelengths that pass a point per unit time [Hz].

c – speed of light, λ – wavelength, v – frequency.

The longer the wavelength, the lower the frequency; the shorter the wavelength, the higher the frequency.

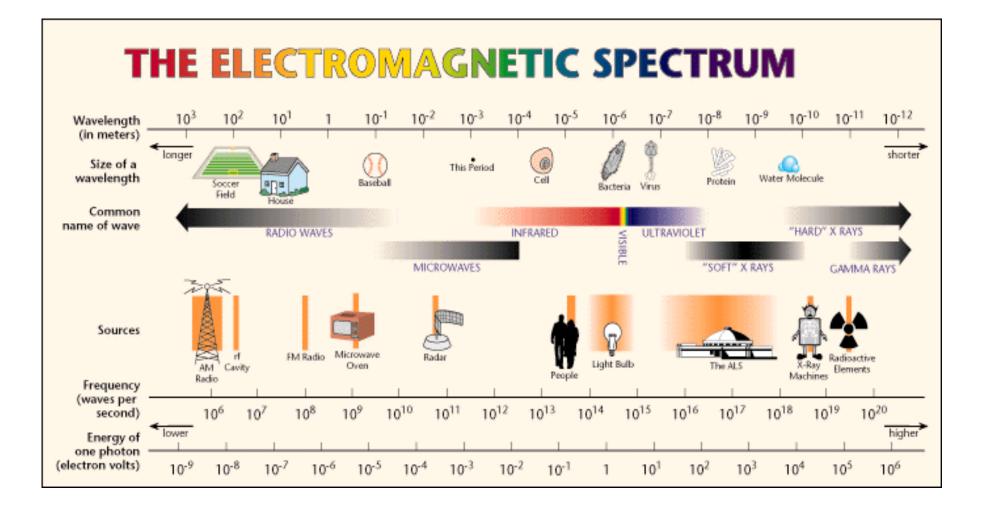
The Electromagnetic Spectrum



Blue: 400–500 nm Green: 500-600 nm Red: 600-700 nm

Optical remote sensing: visible + infrared

Infrared - near infrared (700-1300 nm) + short-waveinfrared (1300 – 2500 nm)





Visible wavelength



Near-infrared wavelength

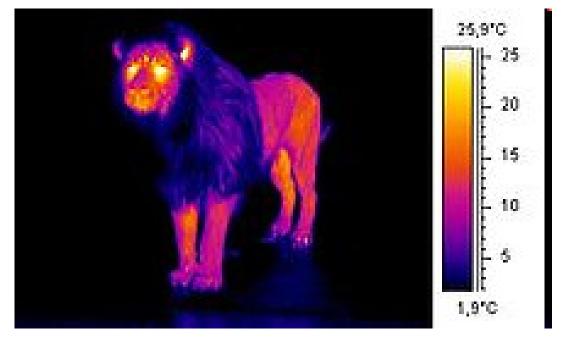


Photo of a lion taken in the mid-infrared (colors contractual)

Himalyan balsam (policeman's helmet) - Impatiens glandulifera

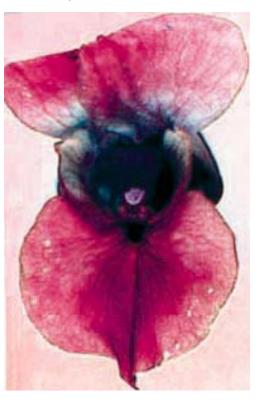
Human Vision



Bee Vision Simulation



Butterfly Vision Simulation



Remote Sensing – electromagnetic I gensen 2000 radiation

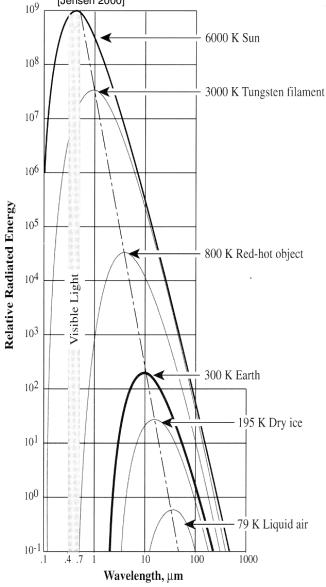
Wave model - James C. Maxwell

Wiens displacement law

(dominant wavelength)

$$\lambda_{\rm max} = k/7$$

k - a constant equaling 2898 μm K T – absolute temp. in degrees K



Blackbody radiation curves

The particle model

Photon energy units:

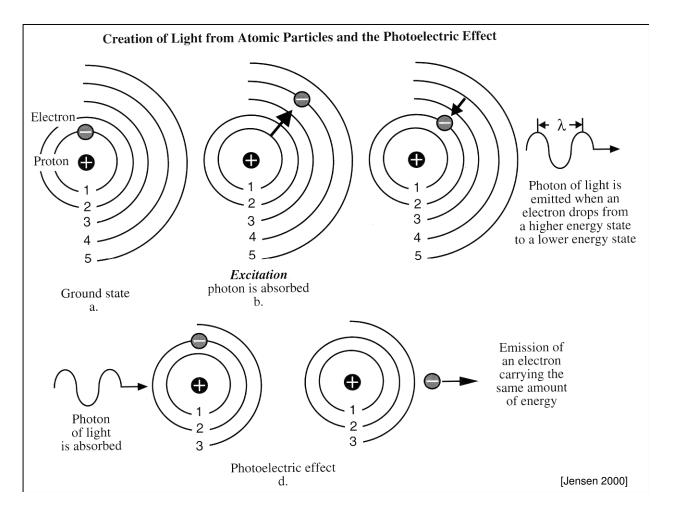
Joules (J) electron volts (eV)

Q = hv

Q - the energy of a quantum measured in Joules (J), **h** - the Planck constant (6.626 x 10^{-34} J s), **v** - the frequency of the radiation,

c – speed of light.

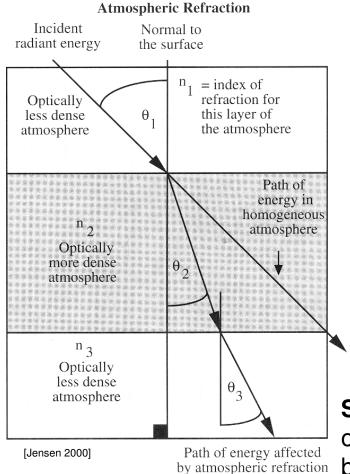
$$\lambda = \frac{hc}{Q}$$
 or $Q = \frac{hc}{\lambda}$



Radiant energy (Q)

- Electromagnetic radiation can be described using both wavelength λ, frequency v and size of radiant energy Q, which causes photochemical reaction in the emulsion film or inducing light-sensitive element (eg CCD).
- Radiant energy Q is understood as the ability of radiation to perform physical work, warm-up or cause changes in state of matter, measured in J or eV.

Energy – Matter interaction in the Atmosphere



Radiant energy - the capacity of radiation within a spectral band to do work

<u>Refraction</u> - the bending of light when it passes from one medium to another

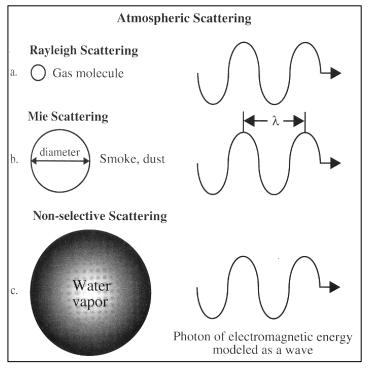
Index of refraction - measure of the optical density of a substance. Atmosphere - 1.0002926 Water - 1.33

 $n = \frac{C}{C_n}$ where: c - speed of light in a vacuum, c_n - speed of light in a given substance.

Snell's law - for a given frequency of light the product of the index of refraction and the sine of the angle between the ray and a line normal to the interface is constant

Energy – Matter interaction in the Atmosphere

Scattering is a general physical process where some forms of radiation, such as light are forced to deviate from a straight trajectory by one or more localized non-uniformities in the medium through which they pass. **The direction associated with scattering is unpredictable**. The direction of **reflection** is **predictable**



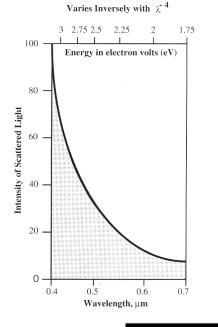
Rayleigh scattering (molecular scattering) – at particles with the effective diameter many times smaller (usually < 0.1) than the wavelength of the incident electromagnetic radiation

UV light at 300 nm is scattered approximately 16 times more than red light at 600 nm

Blue light at 400 um is scattered about 5 times that of red light at 600 nm

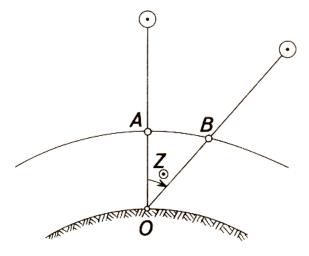


Energy – Matter interaction in the Atmosphere

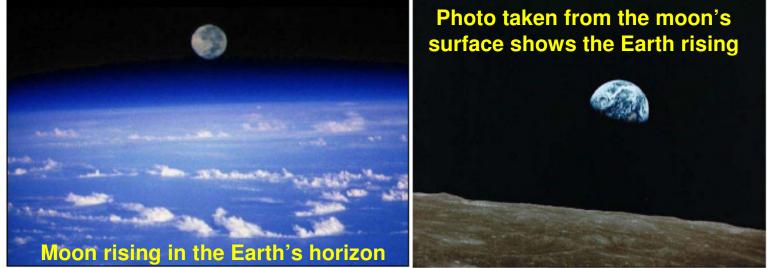


Intensity of Rayleigh Scattering

Rayleigh scattering



 $OA = OB \ sec \ Z_{\odot}$



Energy – Matter interaction in the Atmosphere

Mie scattering (nonmolecular scattering) – at particles with diameters approximately equal to the size of the wavelength of the incident energy

Number of dust particles in 1 cm ³ over:	
Dcean	70 -340
liede vereteine	400 0.000

400 - 8 000
32 000
130 000
170 – 470 000

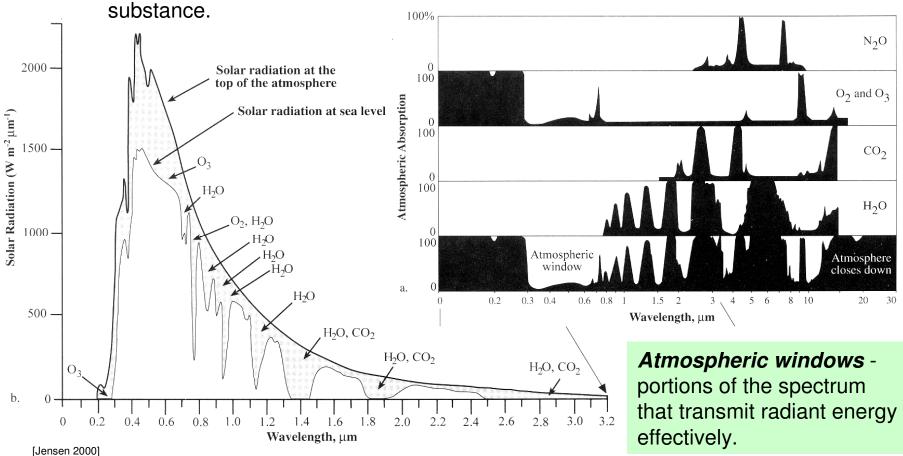
О

Non-selective scattering takes place in the lowest portions of the atmosphere where there are particles greater than 10 times the wavelength of the incident electromagnetic radiation

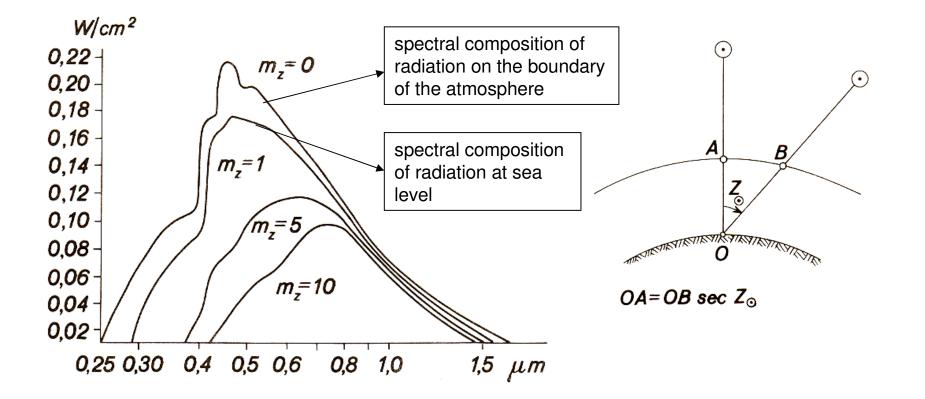
Energy – Matter interaction in the Atmosphere

Absorption is the process by which radiant energy is absorbed and converted into other forms of energy.

An *absorption band* is a range of wavelengths (or frequencies) in the electromagnetic spectrum within which radiant energy is absorbed by a



Energy – Matter interaction in the Atmosphere

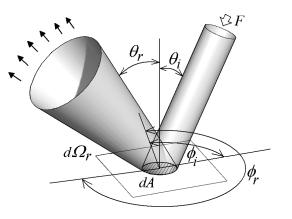


The spectral composition and intensity of solar radiation depending on the thickness of the atmosphere, through which the radiation passes (A.S. Kuczko) [Wójcik]

Radiant flux - *t*he amount of radiant energy onto, off of, or through a surface per unit time and is measured in Watts (W).

Radiation budget equation:

$$\Phi_{i_{\lambda}} = r_{\lambda} + \tau_{\lambda} + \alpha_{\lambda}$$



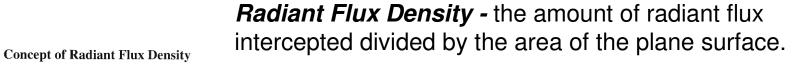
Hemispherical reflectance - the ratio of the radiant flux reflected from a surface to the radiant flux incident to it (dimensionless).

Reflectance factor:
$$\rho_{r_{\lambda}} = \frac{\Phi_{reflected}}{\Phi_{i_{\lambda}}} \bullet 100$$

Hemispherical transmittance - the ratio of the radiant flux transmitted through a surface to the radiant flux incident to it (dimensionless)

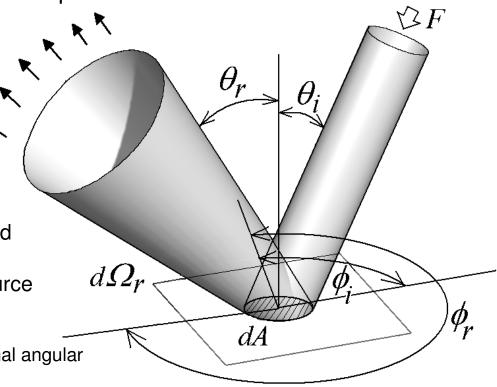
Hemispherical absorptance - the ratio of the radiant flux absorbed by a surface to the radiant flux incident to it (dimensionless)

Hemisphere – half of the sphere



Irradiance - the amount of radiant flux incident per unit area of a plane surface.

Exitance - the amount of radiant flux leaving per unit area of the plane surface.



Radiance - the radiant flux per unit solid angle leaving an extended source in a given direction per unit of projected source area in that direction [W*m^{-2*}sr⁻¹].

Radiant flux, Φ_{2}

Area, A

Radiant flux, Φ_{λ}

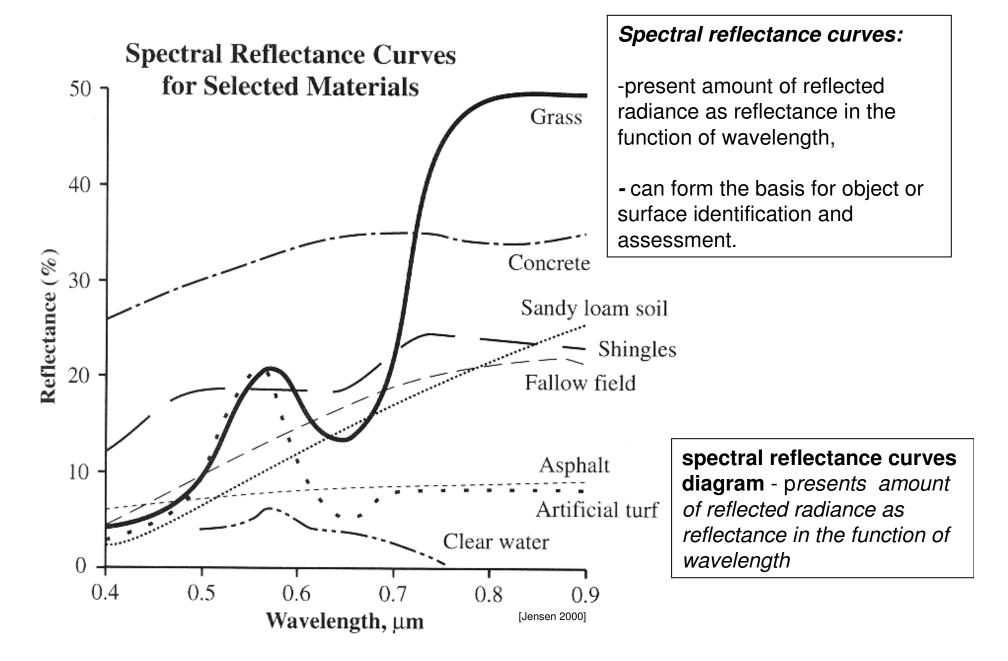
Area, A

Steradian is used to describe two-dimensional angular spans in three-dimensional space.

Irradiance E_{λ} = $\frac{\Phi_{\lambda}}{\Lambda}$

Exitance

 $M_{\lambda} =$



Reflectivity and reflectance refer to the fraction of incident electromagnetic power that is reflected at an interface

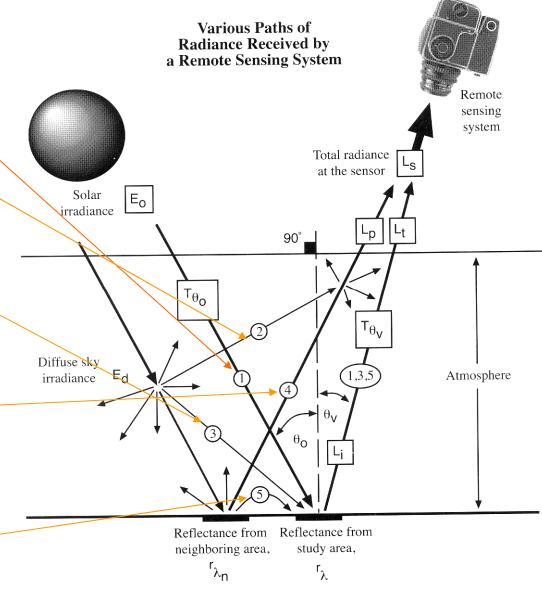
Specular Versus Diffuse Reflectance **Reflection -** fundamental Angle of Angle of characteristics : Angle of Angle of Incidence Exitance Incidence Exitance - the incident radiation, the reflected radiation, and a vertical to the surface from which the angles of incidence and reflection are smooth water measured all lie in the same plane, Near-Perfect Perfect Specular Reflector Specular Reflector - the angle of incidence and the angle of reflection (exitance) are approximately equal. Lambertian surface - the radiant flux leaving the surface is constant for any angle of reflectance to the surface C Near-Perfect Perfect Diffuse Reflector

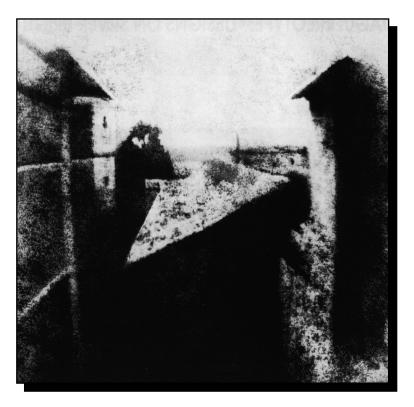
Diffuse Reflector

Lambertian Surface

[Jensen 2000]

Path 1 - contains spectral solar irradiance (Eo) that was attenuated very little Bath 2 anntainating apteal distuan skyring then the (VEO) that never even reaches the Earth's surface (the study **Path** 3 contains energy from area) because of scattering the Sun that has undergone some scattering, absorption and reemission before ip and a standard and a standard and a standard and a standard a sta that was reflected or scattered by nearby terrain (rA) covered by snow, concrete, soil, water, and/or Path 5 is energy that was also the sensor system, reflected from nearby terrain into the atmosphere, but then scattered or reflected onto the study area.



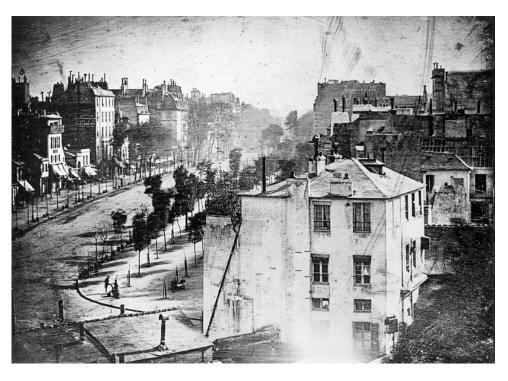


Nicephore Niepce's earliest surviving photograph of a scene from nature, circa 1826, "View from the window at Le Gras" Saint-Loup-de-Varennes (France).

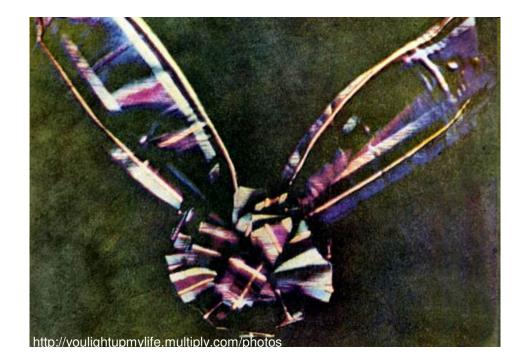
(copyright Gernsheim Collection, Harry Ransom Humanities Research Center, University of Texas)

-Metal plate.

 White bitumen of Judea (derivative of petroleum) which hardens with exposure of light.
 Silver compounds (silver nitrate AgNO₃)



"Boulevard du Temple", taken by Louis Daguerrein late 1838 or early 1839, was the first-ever photograph of people.



James Clerk Maxwell - First Color Photograph - 1861

Balloons





Gaspard Felix Tournachon (*Nadar*) 1858

The oldest surviving aerial photo of Boston in 1860, taken by James Wallace Black, from a balloon.

In 1858, Nadar applied for a patent for an "*aerial survey* - the mapping of the land from a series of overlapping aerial photographs "

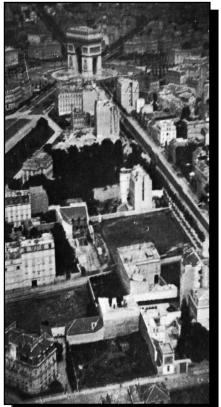


A portion of an aerial photograph of downtown Boston, MA. obtained by Black and King at altitude of 400 m. First aerial photograph taken from a captive ballon in the USA. Balloons



Balloon Intrepid tethered during the Civil War battle of Fair Oaks on June 1, 1862

(copyright Smithsonian Institution, Washington, D.C.).



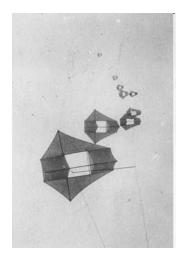
Oblique photograph obtained from the *Hippodrome* Balloon using a multiband camera (1868). *Nadar*

Balloons



Paris, the pont Louis-Phillippe

Gaston Tissandier and Jacques Ducom (dry-plate process)



a train of nine Conyne kites

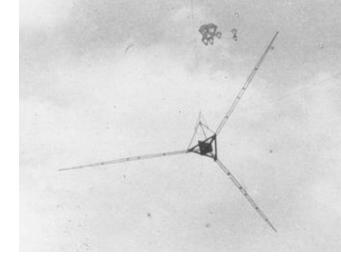
Kites

George Lawrence - Captive Airship



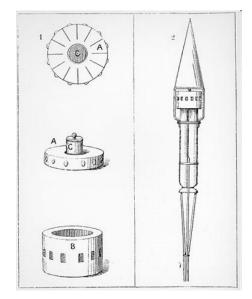
Six weeks after the April 18, 1906





the kite train and camera-steadying mechanism 'Captive Airship."

1908



"Photo rocket" conceived by the Frenchman Amedee Denisse in 1888



Rockets



View of a village in Sweden thought to be taken by Alfred Nobel's photo rocket in 1897

German landscape photographed in 1904 by the camera system of Alfred Maul, a pioneer in photo rocketry.

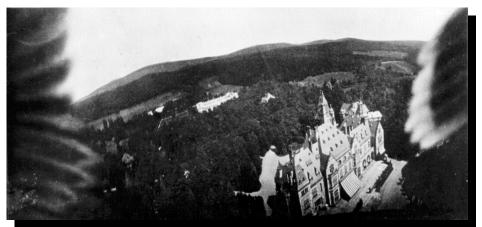
Pigeons



Deutsches Museum, Munich

Pigeon with miniature camera Patented in 1903





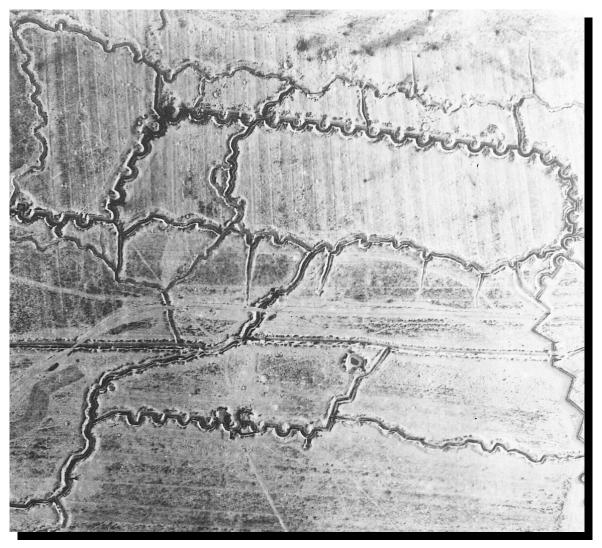
Aeroplanes



The Wright Flyer built by Orville and Wright. The first successful heavier-than-air engine-powered flight on December 1903

11 seconds and 40 meters

World War I



"a most disgraceful thing to have attempted"

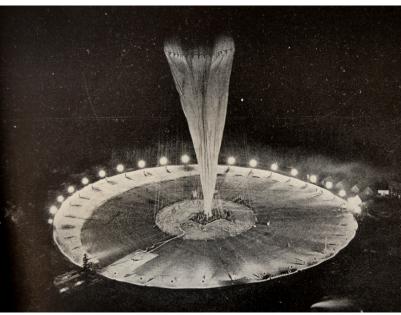
Vertical photography of World War I trenches in Europe.



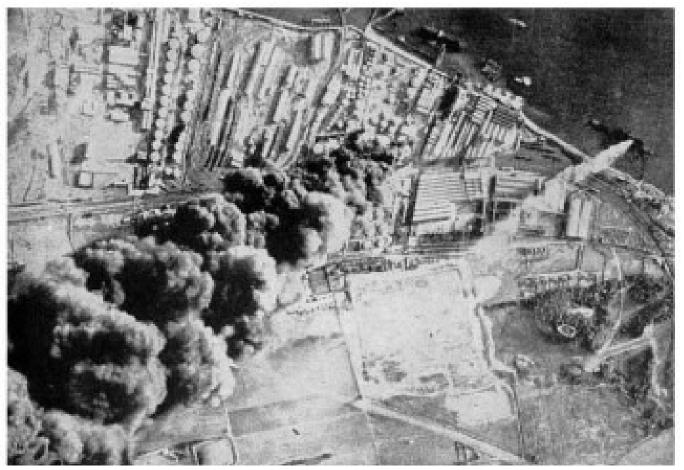
World War I

An aerial photographer and pilot in a Curtiss AH-13 airplane with Graflex camera in 1915





Before 1946, the highest pictures ever taken of the Earth's surface were from the helium-filled Explorer II balloon, which had ascended 13.7 miles in 1935, high enough to discern the curvature of the Earth.



A German photograph of a direct hit on the oil storage tanks at Purfleet, near <u>Tilbury</u> on the <u>Thames</u>

http://nzetc.victoria.ac.nz/tm/scholarly

World War II



An aerial reconnaissance photograph of German invasion barges



German invasion barges in Boulogne Harbour, France, June 1940.

World War II



Craters surround a site at Peenemunde in Mecklenburg-Vorpommem, Germany on September 2, 1944, following an Allied bombing raid on the site where the V weapons were designed and tested



A German photograph of a direct hit on the oil storage tanks at Purfleet, near Tilbury on the Thames.

World War II

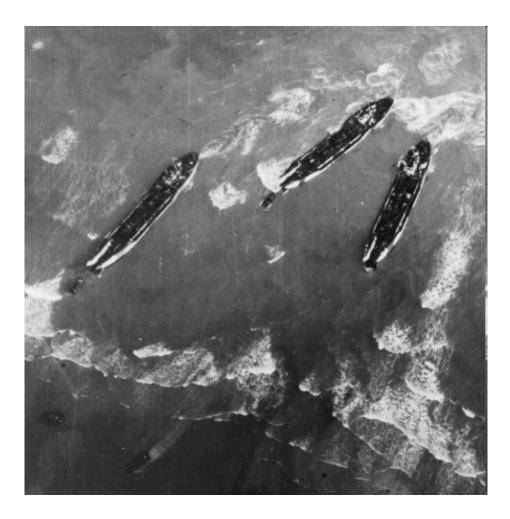




US 8th Air Force B-17 Flying Fortress over Berlin, Germany in World War II.







Operation Market Garden in September 1944

D-Day on June 6, 1944 of the Allied invasion

Warsaw ghetto

Stuttgart



http://gizmodo.com





Project GENETRIX Balloon during launch

Cold War

Grayback Program Project GENETRIX Balloon



This image, taken through a telescope, is of the newly designed super pressure balloon at float altitude over Antarctica.

Credit: Columbia Scientific Balloon Facility

Cold War



Lockheed U-2 high altitude reconnaissance aircraft 1956

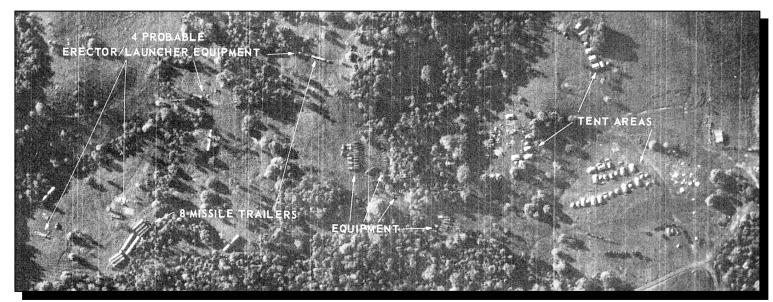


Francis Gary Powers

Two high resolution cameras:

long-focal-length spotting camera (resolution 60 cm to 1m from 20 000 m)
tracking camera to record the flight path

Cold War

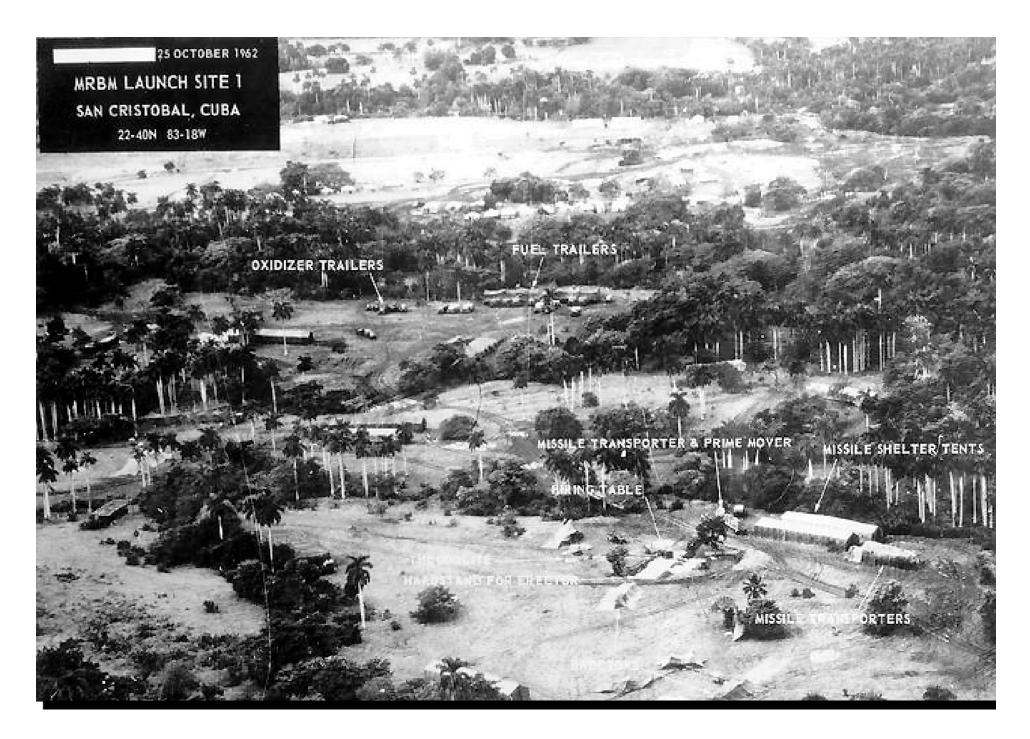


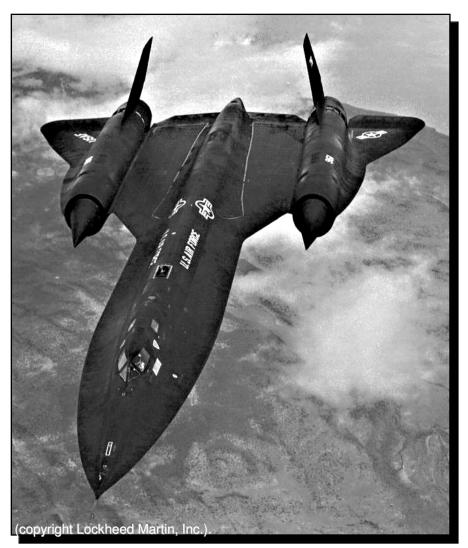
[Jensen 2000]

A portion of the U-2 aerial photograph of San Cristobal, Cuba, 1962 obtained at an altitude of 21 000 m.

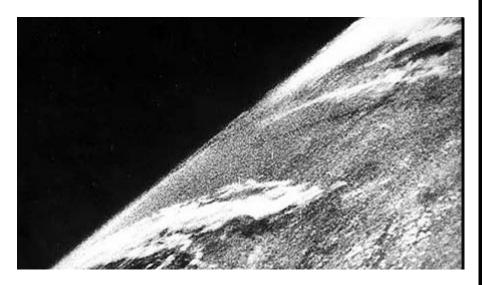
Low-oblique aerial photograph of ballistic missile launch site obtained by low-flying RF-101 aircraft.







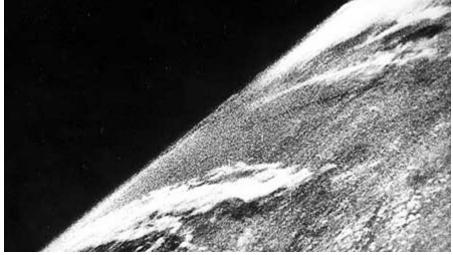
Lockheed SR-71 reconnaissance aircraft



First Image From Space - 1946



First Image Of The Entire Earth - 1968



24 October 1946 White Sands (New Mexico)

Rocket V2

7 March 1947 Gulf of California



V-2 ROCKET-EYE VIEW FROM 60 MILES UP



ROCKET FIRED AT WHITE SANDS PROVING GROUND, JULY 26,1948

8- BLACK RANGE

DISTANCE FROM CAMERA TO HORIZON-700 MILES

13- SANDIA MTS.

- 14- VALLE GRANDE MTS.
- 15- RIO GRANDE

11- MT. TAYLOR

16- SANGRE DE CRISTO RANGE

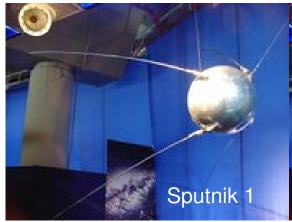
AREA SHOWN APPROXIMATELY 800.000 SO.MILES

DISTANCE ALONG HORIZON-2700 MILES

INSTRUMENTATION AND PHOTOGRAPHY BY APPLIED PHYSICS LABORATORY OF THE JOHNS HOPKINS UNIVERSITY FOR THE BUREAU OF ORDNANCE, U.S.NAVY

A panorama from a July 1948 V-2 shot here.)

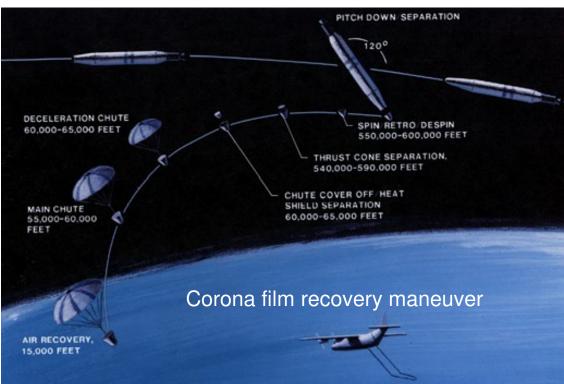
5- GILA RIVER



Corona sensors: KH-1, KH-2 ... KH-6

Spatial resolution:

KH-1: 12 m (1960) KH-4B: 1.4 – 1.8 m (1972)





Mys Shmidta Air Field, Chukotka U.S.S.R. August 18, 1960. First recovered reconnaissance image from the Corona spy satellite



The first picture of the whole Earth - 1968



Satellite image which, according to Americans, confirms that North Korea is preparing to launch a ballistic missile 30.03.2012